

# Supergiants and AGB stars at high angular resolution

**Pierre Kervella**

CNRS/U. de Chile UMI FCA 3386 & LESIA, Paris Observatory  
+ many collaborators: M. Montargès, W. Homan, A. Richards,  
E. Lagadec, L. Decin, K. Ohnaka, E. O’Gorman,...

*photo: S. Guisard*

# RSGs and AGBs

- Red supergiants and asymptotic giant branch stars experience intense mass loss
- Extraction of material from the star, formation of molecules and dust are complex
- High angular resolution techniques now resolve the key spatial scales for nearby stars
- Examples of recent results on Betelgeuse and L<sub>2</sub> Pup (+ see other presentations !)

# Betelgeuse

Wind acceleration

0.025 arcsec

0.25 arcsec

2.5 arcsec

4 arcmin

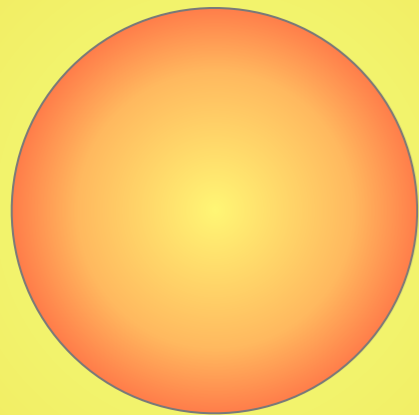
Internal envelope  
1-10  $R_*$

Intermediate envelope  
10-100  $R_*$

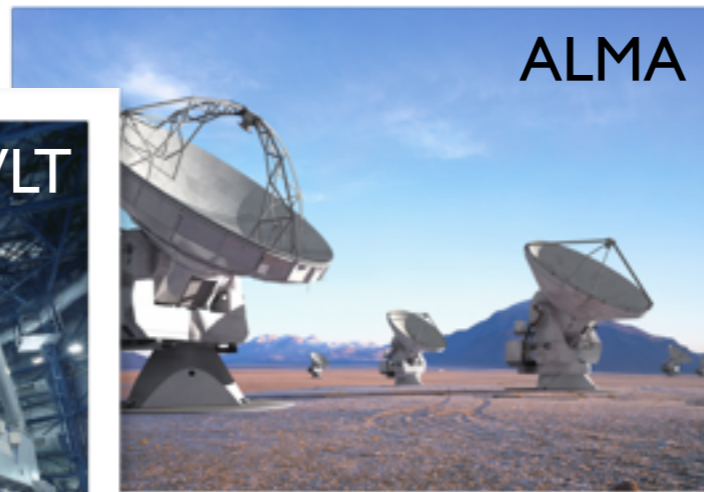
External envelope  
100-10<sup>4</sup>  $R_*$

MIS

Photosphere



VLT



ALMA



VLTI

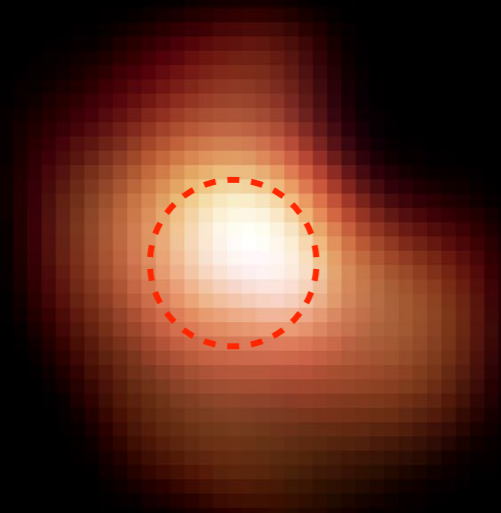


HST

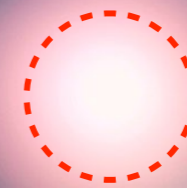


Herschel

ZIMPOL (0.5-0.7 $\mu\text{m}$ ) 2015

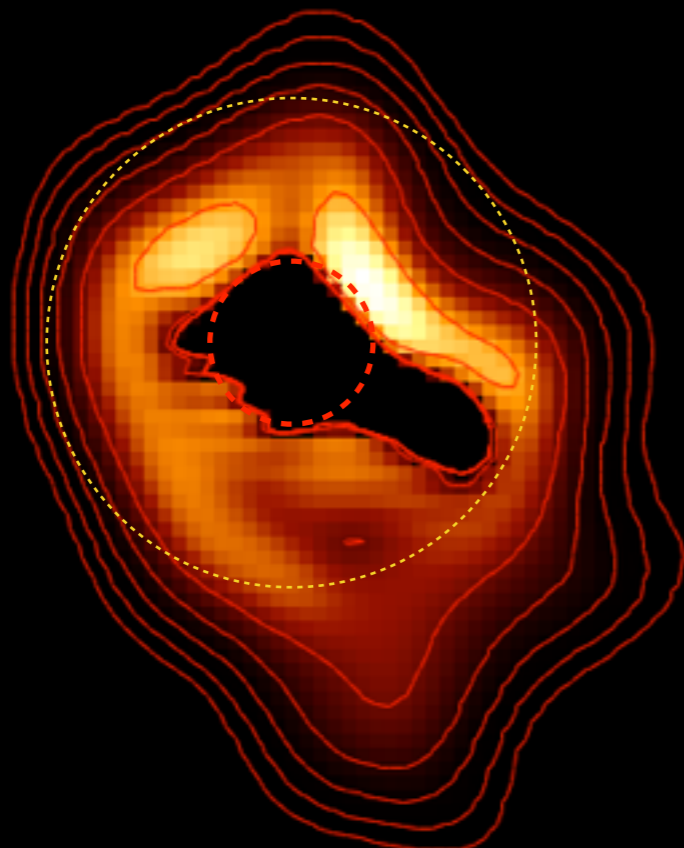


NACO (1.0-2.2 $\mu\text{m}$ ) 2009

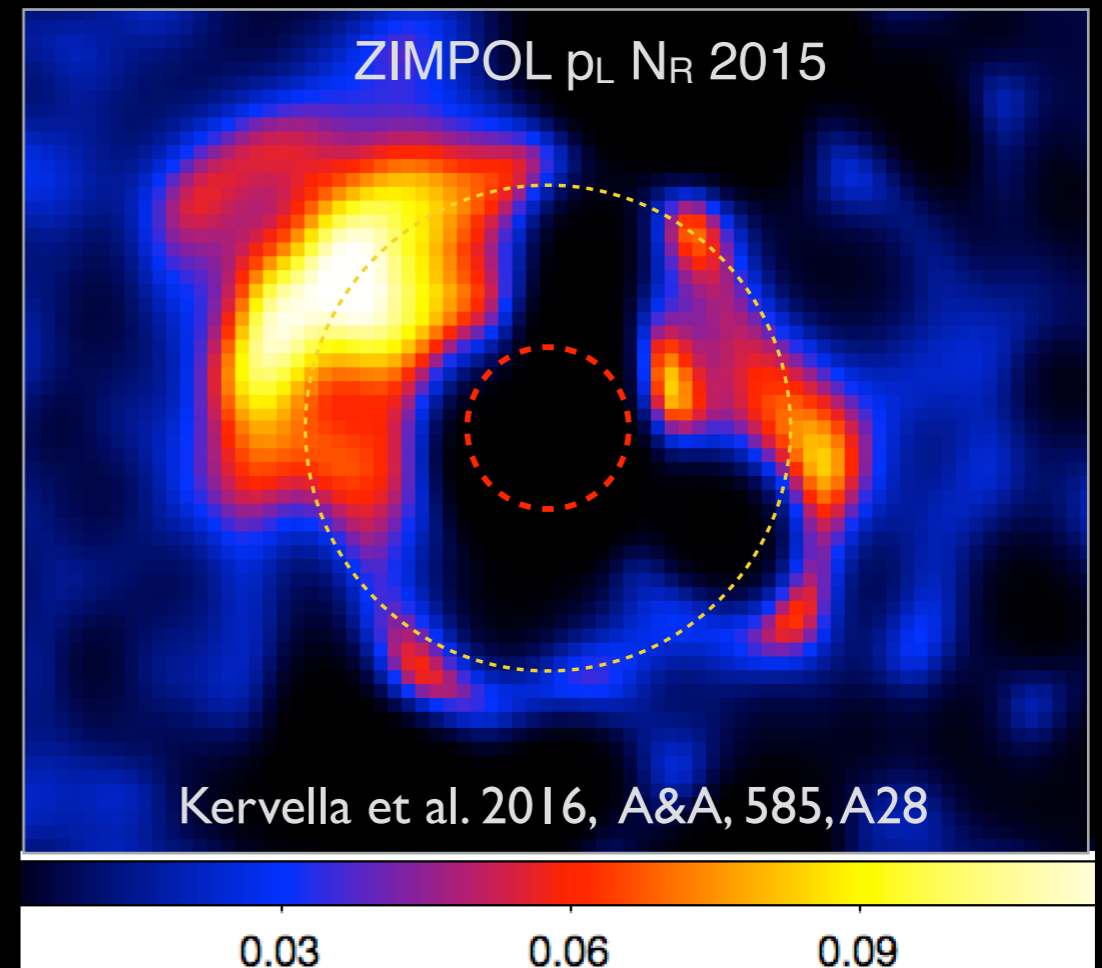


Kervella et al. 2009, A&A, 504, 115

ZIMPOL H $\alpha$  2015



ZIMPOL p<sub>L</sub> N<sub>R</sub> 2015



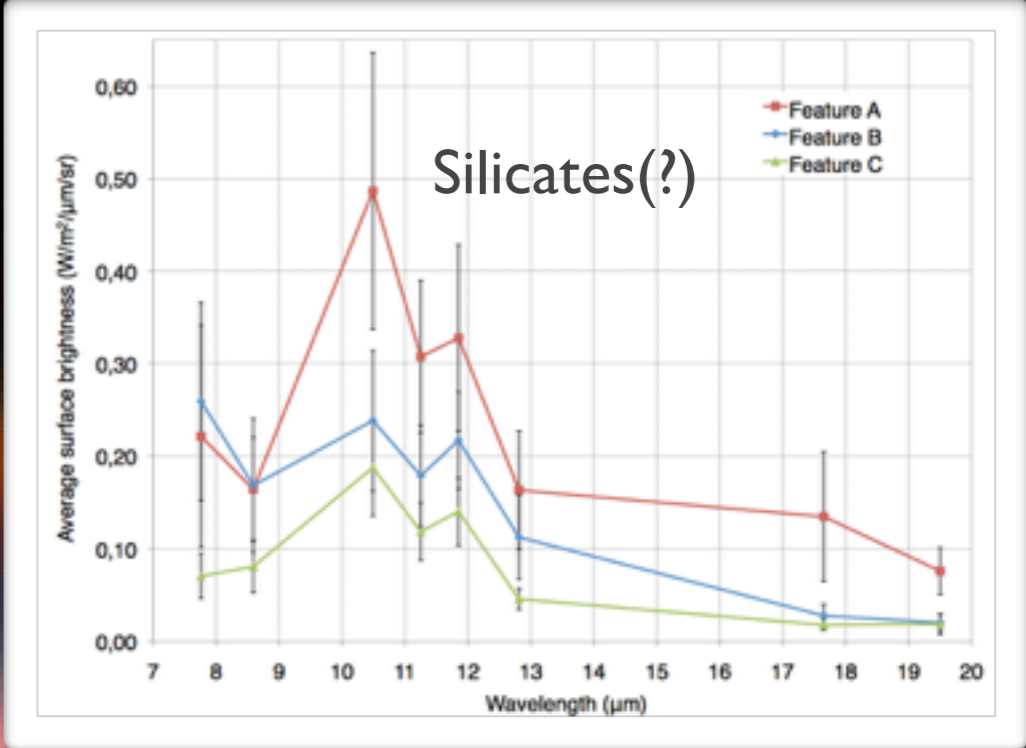
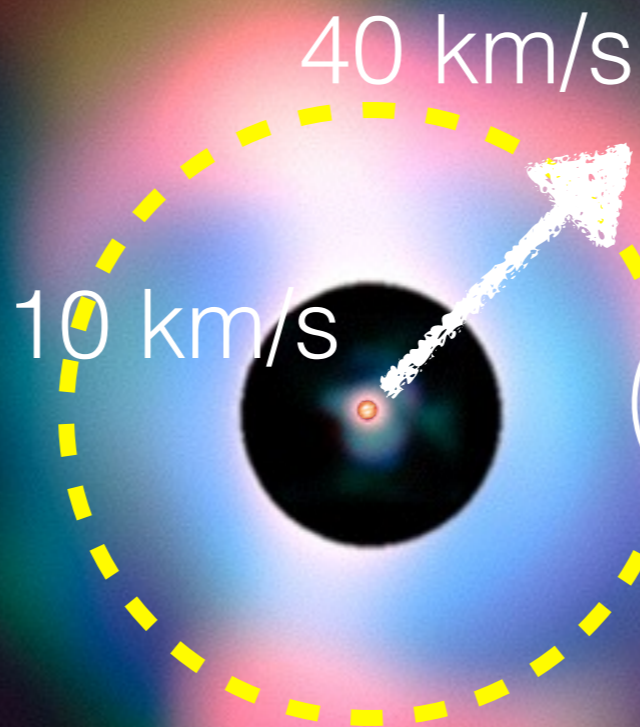
Kervella et al. 2016, A&A, 585, A28

0.03

0.06

0.09

Dust shell



1"

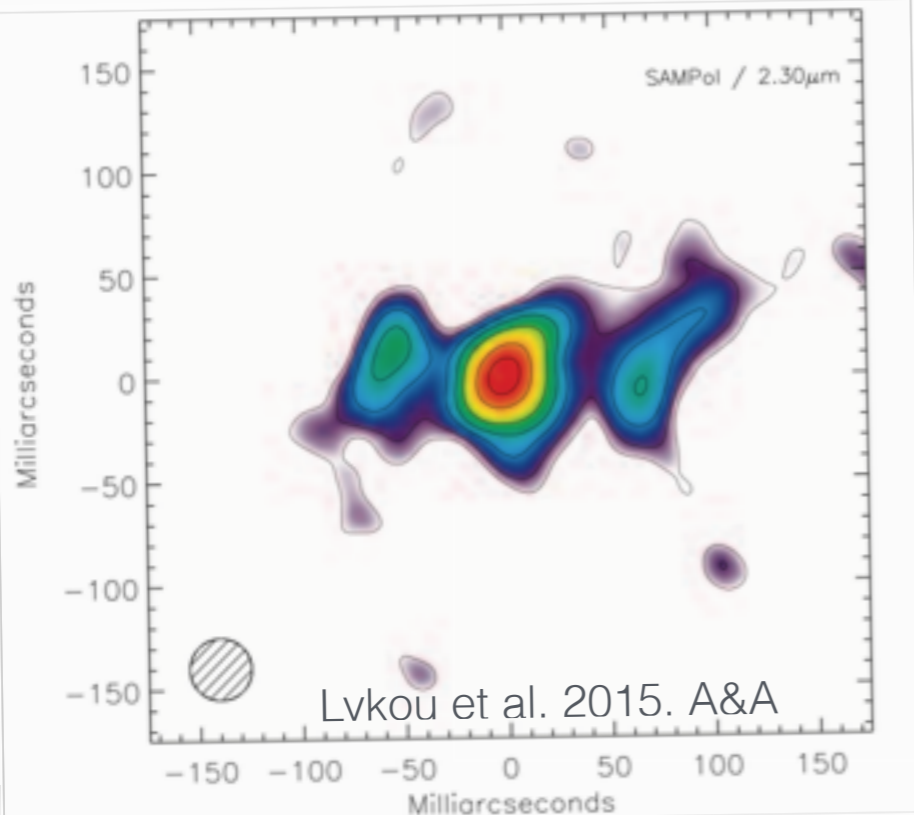
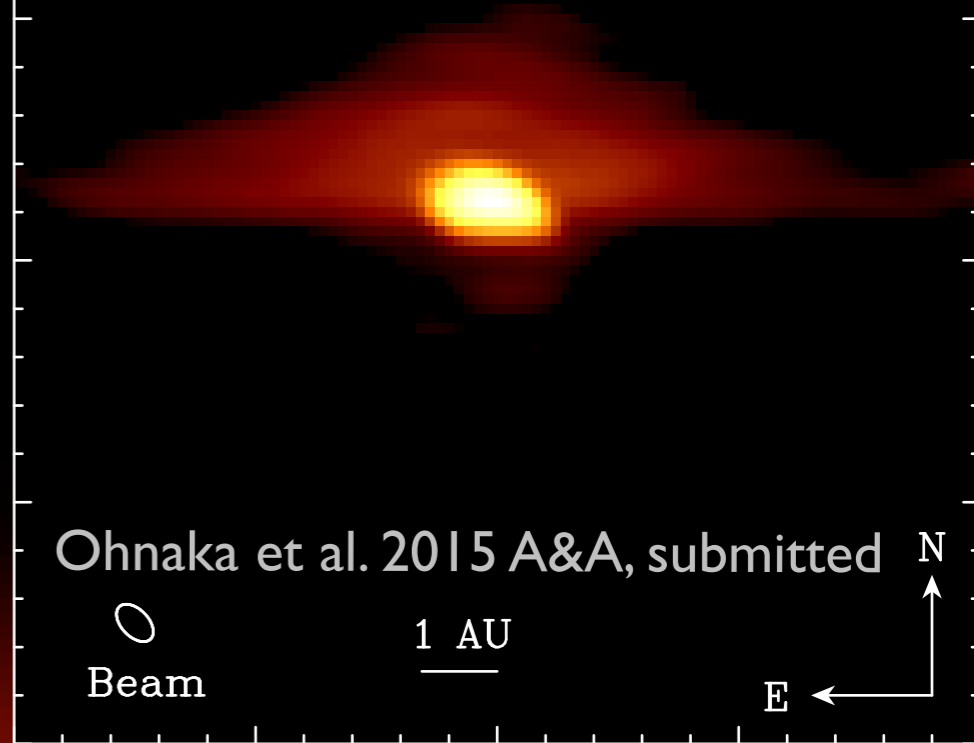
VISIR (2011)  
8 - 20  $\mu m$

# L2 Puppis

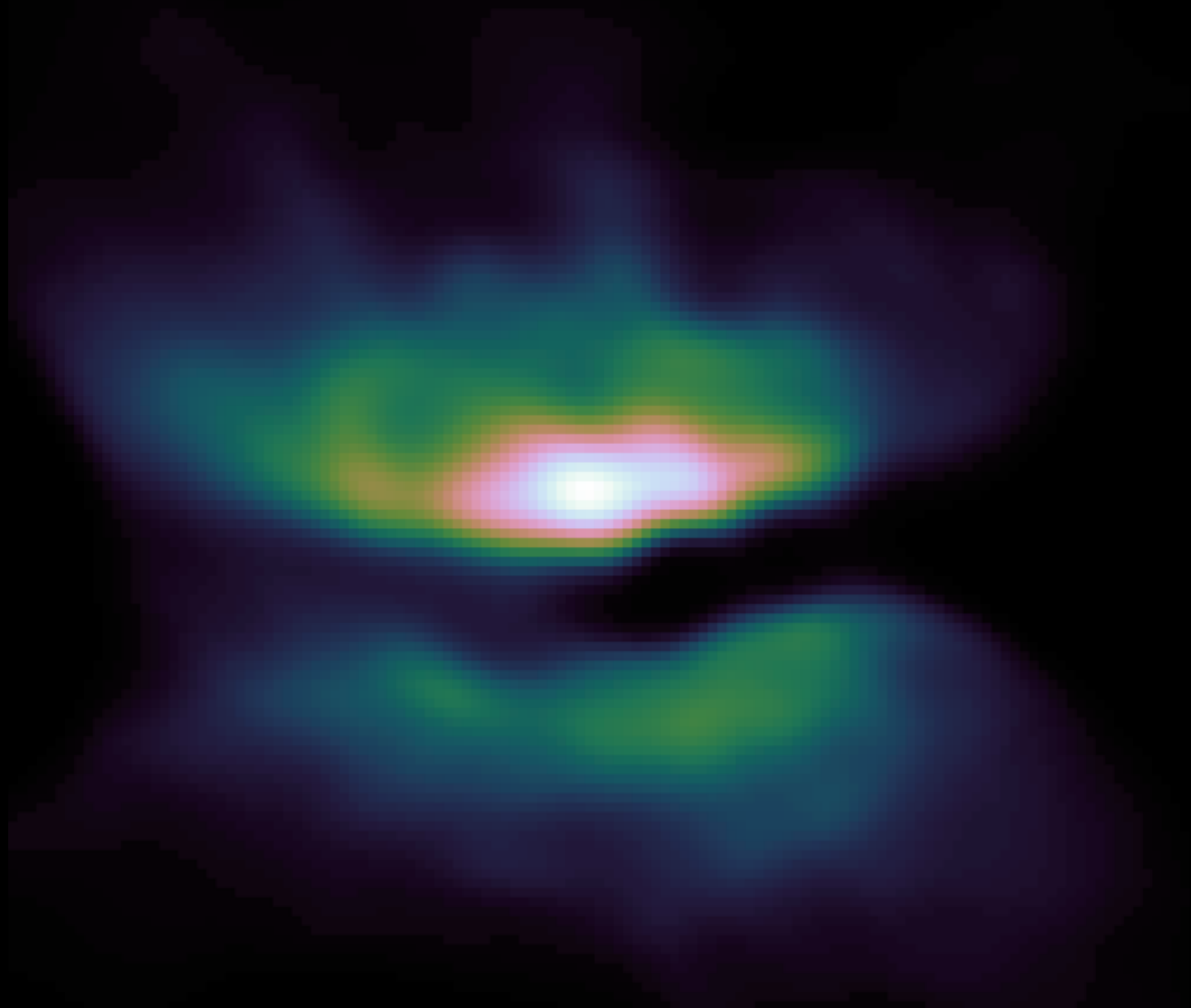
- M5III, SRa (Mira-like) star with  $P=141$  d
- Mass : from 0.7  $M_{\text{sun}}$  (Lykou et al. 2015) up to 2  $M_{\text{sun}}$  (Kervella et al. 2014)
- Suspected binary star from Hipparcos astrometric wobble)
- Second nearest AGB star ( $64 \pm 4$  pc,  $m_v \sim 5$ ) after R Doradus ( $\sim 55$  pc)
- Long term, slow dimming event ongoing since 15 years (Bedding et al. 2002)

2.17  $\mu\text{m}$

b. NACO speckle + AMBER

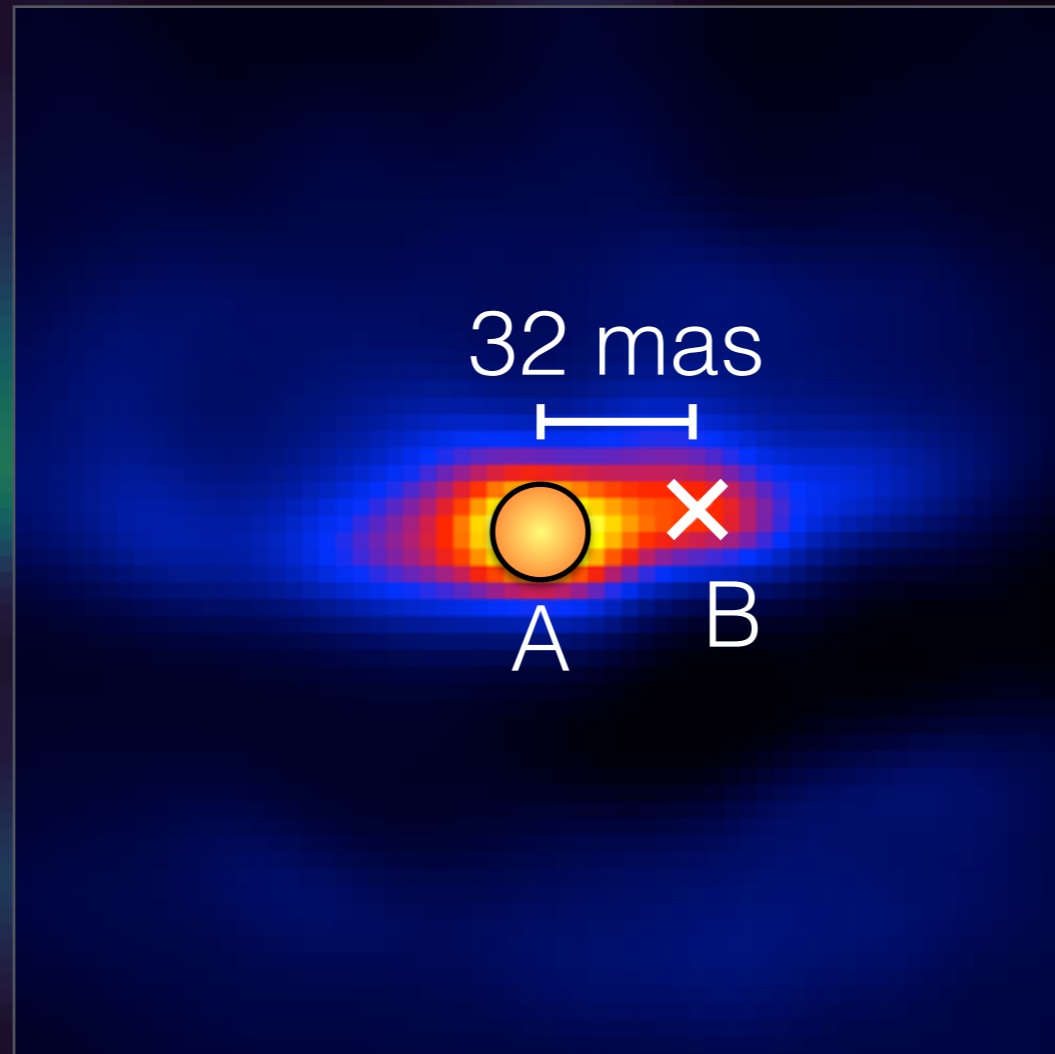


VLT/VINCI



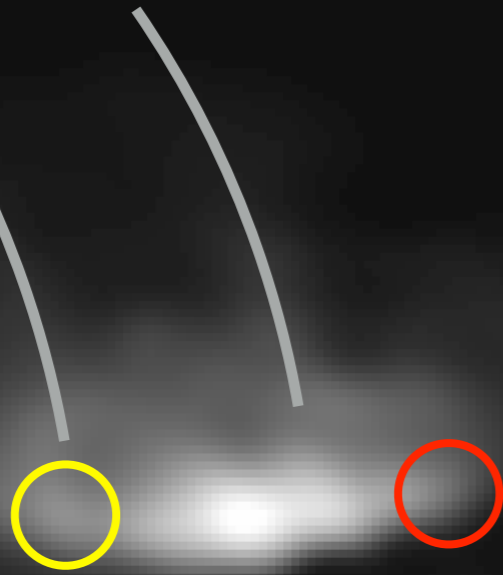
0.65  $\mu\text{m}$  (SPHERE/ZIMPOL)





# Degree of linear polarization

$N_R$  band (646 nm)

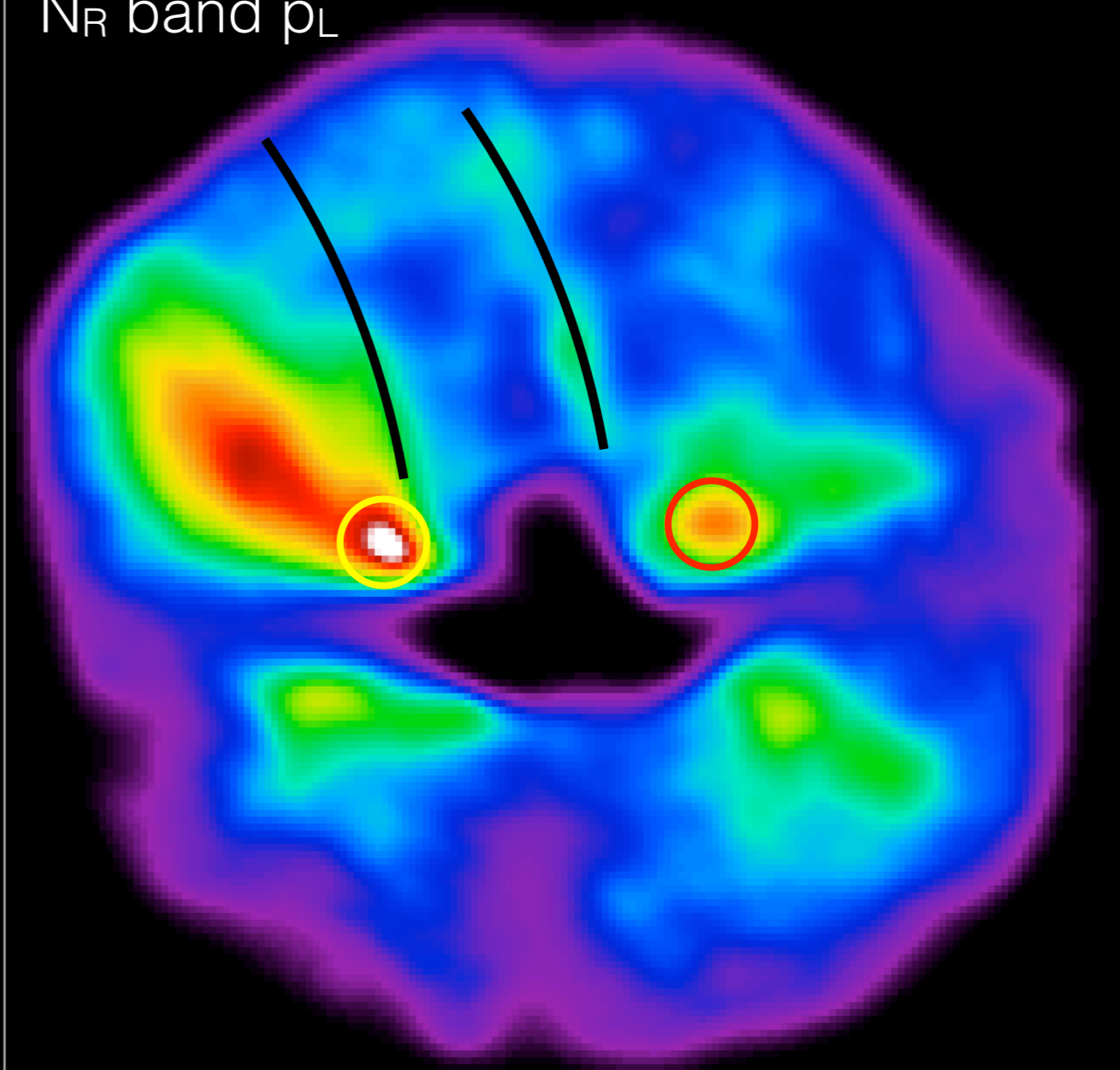


1080

15400

158000

$N_R$  band  $p_L$



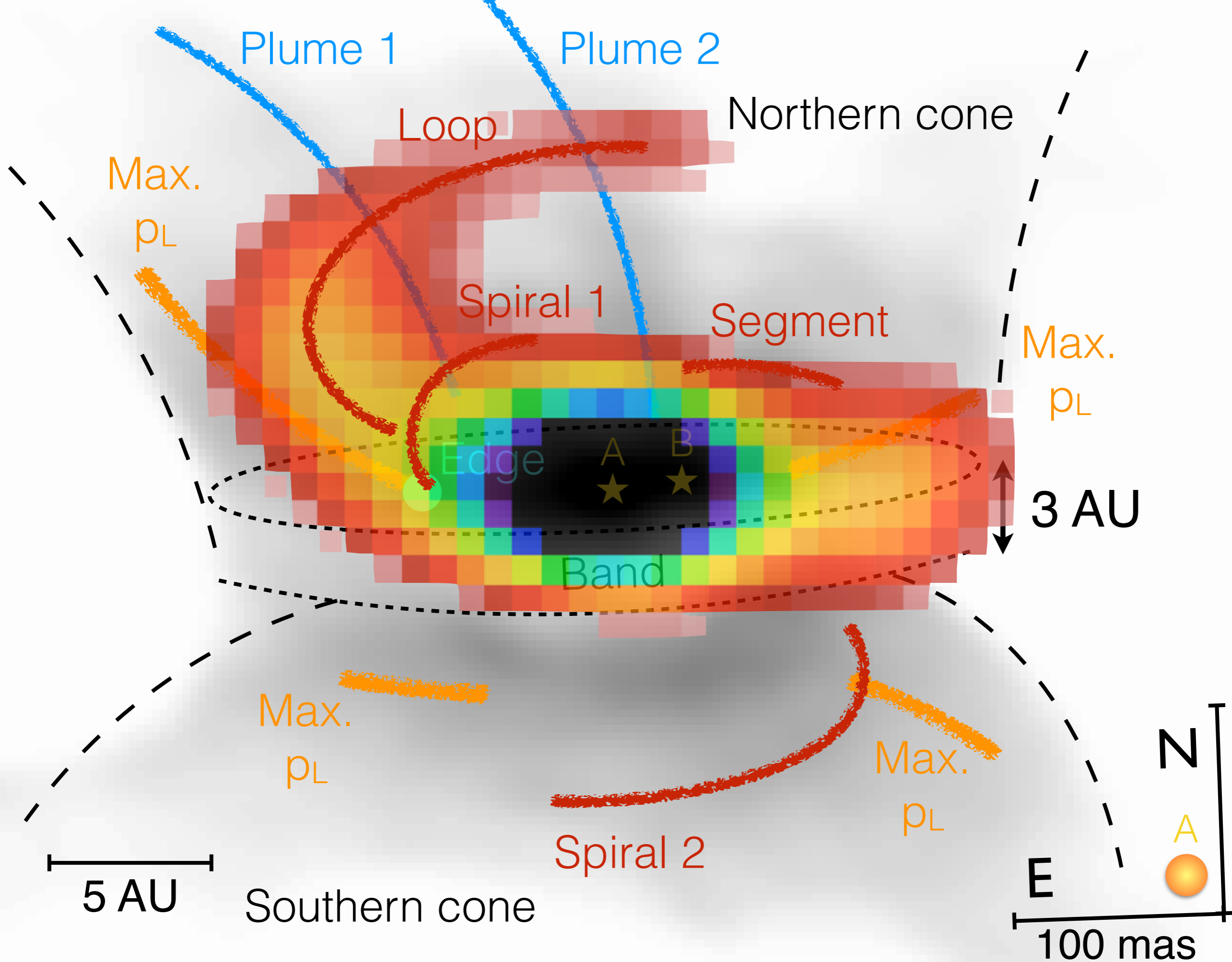
0.1

0.2

0.3

0.4


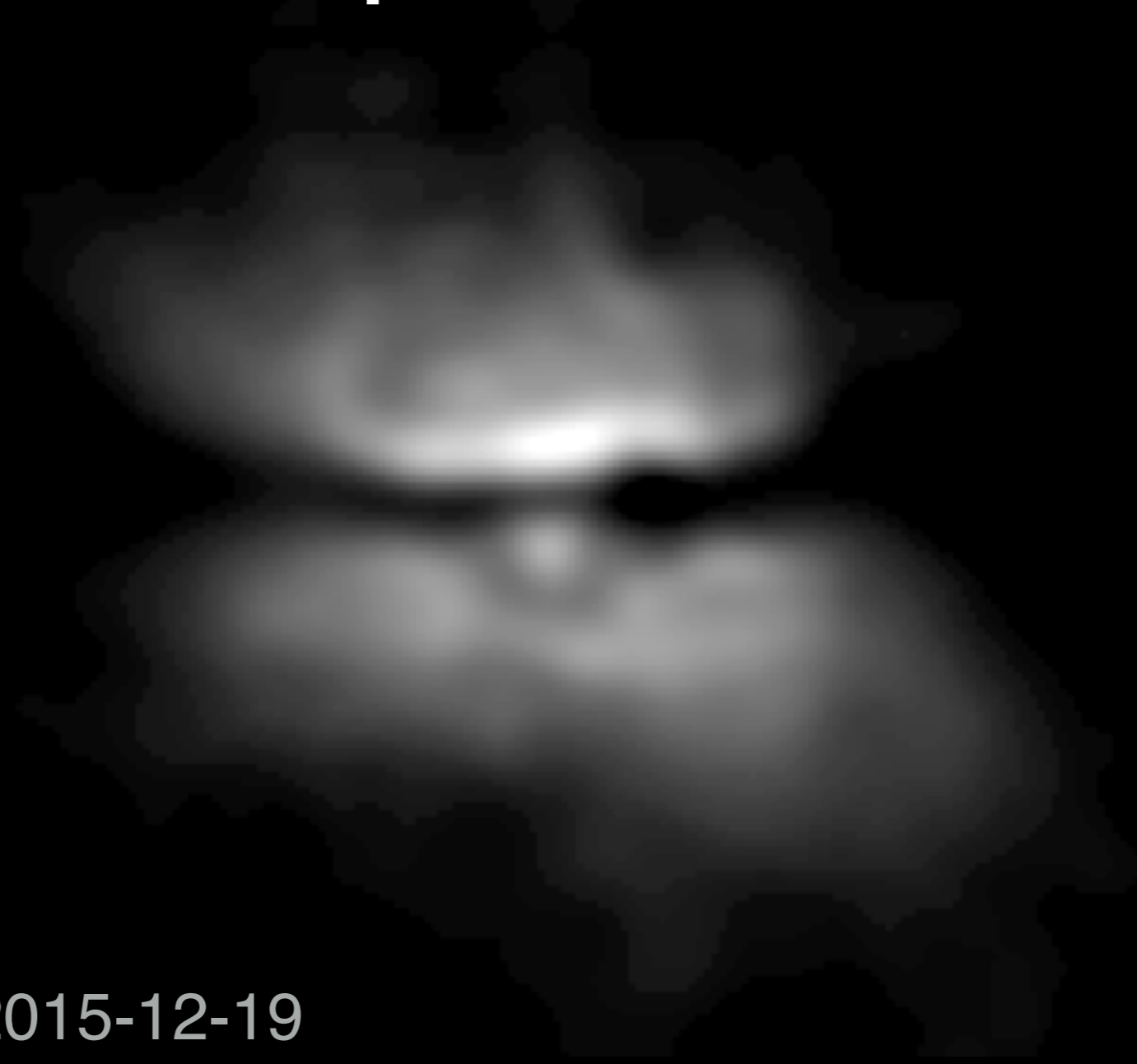
0.5



# L2 Pup evolution

2015-12-19

0.1''

A horizontal scale bar with vertical end caps, indicating a length of 0.1 arcseconds.



- What creates the plumes ?
- How does the AGB wind interact with the disk ?
- How do the streamers form ?
- What is the nature of the companion ? Is it real ?