

Betelgeuse workshop: linearly polarised spectrum of μ Cep

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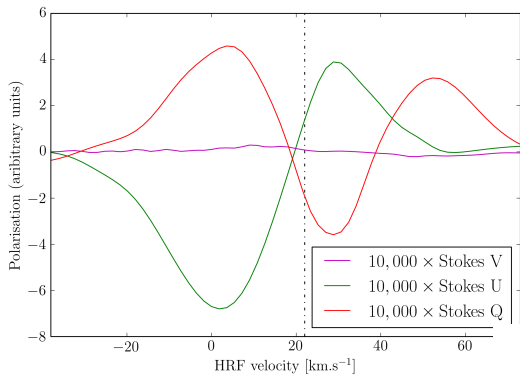
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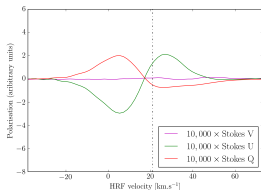
A. Lèbre, J. Morin, A. López Ariste, M. Aurière, Ph. Mathias, E. Josselin, P. Petit

Paris - September 2016

Typical polarisation scales for μ Cep



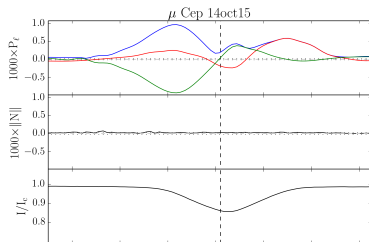
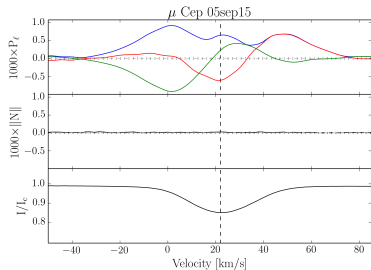
- Ratio $P_{\ell}^{\mu \text{ Cep}} / P_{\ell}^{\text{Betelgeuse}}$
- ~ 2.28
- μ Cep: best target to extend the depolarisation of continuum hypothesis ?



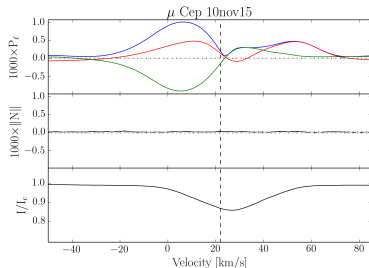
Betelgeuse polarisation scales

1- Linearly polarised spectrum of μ Cep: variability (LSD profiles* view)

*acquired during our Large Program with the Narval instrument



- Like Betelgeuse Variability on months / weeks
- ➔ Consistent with convective time scales



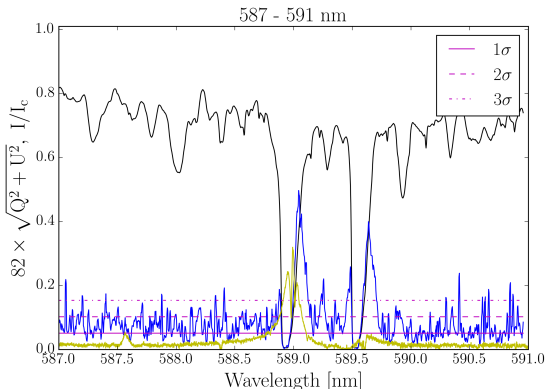
1- Linearly polarised spectrum: Origin (1/2)

Stokes Q & Stokes U

$$P_\ell = \sqrt{Q^2 + U^2}$$

P_ℓ sun

- Non Zeeman origin
- Na D2 amplitude \sim D1 amplitude (consistent with depolarisation of continuum)
- Interpretation for Betelgeuse seems again valid for μ Cep

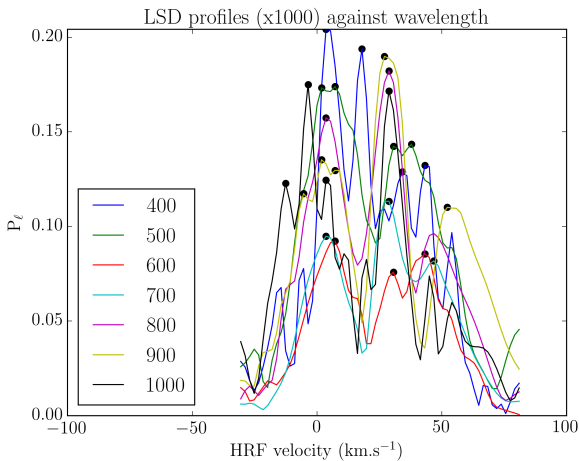


Linearly polarised spectrum of μ Cep around the Na doublet

Polarisation in spectral lines

1- Linearly polarised spectrum: Origin (2/2)

→ Wavelength dependence more complex

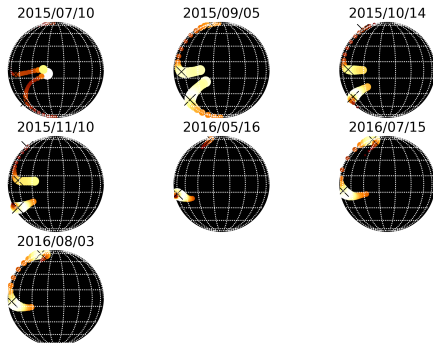


Linear polarisation variation with wavelength.

→ *If we look at the variation of the LSD profile with wavelength at the position of a peak, the decrease with wavelength is not so obvious.*

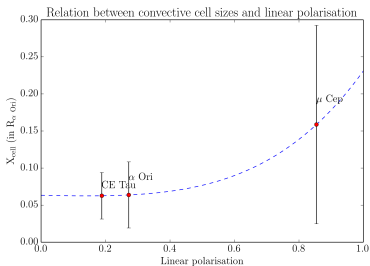
Hot spots model: variation with time

→ Variation with time of the inferred position of "bright spots"



Opened questions: Study the surface dynamics with high resolution spectra

- Best targets to match interferometric observations (μ cep, antares ??)
- Forward modelling of the Q/U signals \rightarrow models for spectropolarimetry
- From the models: constrains the relation $P_\ell \propto x_{\text{cell}}$: infer stellar parameters, continuum polarisation
- Relation with mass loss ?



Stellar parameters from Josselin & Plez 2007 are used and P_ℓ values come from our LP.